# Studying quantum beamstrahlung and nonperturbative QED with beam-beam collisions at FACET-II

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# Motivation: beamstrahlung in future linear colliders



- Why a linear electron-positron collider?
  - Clean interaction (unlike protons  $e^+/e^-$  are elementary particles)
  - Initial state known (protons: parton distribution functions)
  - TeV-scale  $e^+/e^-$  ring not feasible due to synchrotron radiation energy loss
- Beamstrahlung at the interaction point:
  - High luminosity  $\rightarrow$  high charge density  $\rightarrow$  strong fields  $\rightarrow$  beamstrahlung
  - Beamstrahlung theory has never been tested in the quantum regime

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## Motivation: studying nonperturbative QED

#### Important scales of QED

Energy	${\cal E}=mc^2$	$10^6{ m eV}$	relativistic effects
Length	$\lambda_C = \hbar c / (mc^2)$	$10^{-13}\mathrm{m}$	quantum fluctuations
Field strength	$E_{ m cr}=(mc^2)^2/( e \hbar c)$	$10^{18}\mathrm{V/m}$	nonperturbative effects

Electron/positron mass (m) and charge (e < 0) determine fundamental scales

Relativity: Dirac equation	Quantum fluctuations: QFT
- Changed dispersion relation: $\epsilon = mv^2/2$ vs. $\epsilon = \gamma mc^2$	<ul><li>Virtual particles</li><li>Lamb shift of atomic levels</li></ul>
<ul> <li>Spin degree of freedom</li> <li>Antiparticles</li> </ul>	<ul> <li>Anomalous magnetic moment</li> <li>Running coupling constant</li> </ul>
<ul> <li>Antiparticles</li> </ul>	- Running coup

- At each fundamental scale the theory changes qualitatively
- Nature surprised us whenever we tested a fundamental scale
- QED critical field *E*<sub>cr</sub> has never been exceeded experimentally

We use natural units from now on  $\epsilon_0 = \hbar = c = 1$  (often restored for clarity)

# The QED critical field & spontaneous pair production

- According to quantum mechanics (Heisenberg uncertainty principle) the vacuum contains virtual electron-positron pairs (pictorial model)
- Spatial scale of these quantum fluctuations:  $\lambda_C = \hbar/(mc)$
- If an electric field is able to transfers the rest energy  $2mc^2$  to these pairs within their lifetime, they become real:  $E_{\rm cr} = mc^2/(|e|\lambda_C)$



	$\sim \hbar \omega$	Future facilities	<i>I</i> (intensity)	current
optical	$1\mathrm{eV}$	APOLLON, ELI,	$10^{24-25}{ m W/cm^2}$	$10^{22}\mathrm{W/cm^2}$
x-ray	$10\mathrm{keV}$	LCLS-II, XFEL,	$10^{27}{ m W/cm^2}$ (if focused)	$10^{21}\mathrm{W/cm^2}$

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## Reaching the QED critical field in a laboratory experiment

- The critical field cannot be reached directly (with existing technology)
- Fortunately, the electric/magnetic field is not Lorentz invariant:

$$oldsymbol{E}' = \gamma (oldsymbol{E} + eta imes oldsymbol{B}) - rac{\gamma^2}{\gamma + 1}eta (eta oldsymbol{E}), 
onumber \ oldsymbol{B}' = \gamma (oldsymbol{B} - eta imes oldsymbol{E}) - rac{\gamma^2}{\gamma + 1}eta (eta oldsymbol{B})$$

• Decisive measure: electric field in the electron rest frame (E\*):

$$\chi = rac{\sqrt{pF^2p}}{E_{
m cr}mc^2} = rac{E^*}{E_{
m cr}}, \quad E_{
m cr} = rac{m^2c^3}{\hbar\,|e|} pprox 1.3 imes 10^{18}\,{
m V/m}$$

**Electron-laser collisions** 

$$\chi \approx 0.5741 \, \frac{\epsilon}{10 \, \text{GeV}} \, \sqrt{\frac{l}{10^{20} \, \text{W/cm}^2}}$$

*I*: laser intensity  $\epsilon$ : electron energy (head-on collision)

Static magnetic field

$$\chi = \gamma \frac{B}{B_{\rm cr}}, \quad B_{\rm cr} = \frac{m^2 c^2}{(\hbar |e|)}$$

If 
$$\epsilon \gg mc^2$$
 and  $E \ll E_{\rm cr}$ ,  $B \ll B_{\rm cr}$ :  
only  $\chi$  is important

Ritus, J. Sov. Laser Res. 6, 497-617 (1985); Di Piazza et al., Rev. Mod. Phys. 84, 1177 (2012)

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## Probing the QED critical field in beam-beam collisions

Quantum parameter (beamstrahlung parameter)

$$\chi = \frac{|e|}{m^3} \sqrt{p^{\mu} F_{\mu\nu}^2 p^{\nu}}, \quad \chi_{\max} \approx \frac{2N r_e^2 \epsilon}{\alpha m \sigma_z (\sigma_x + \sigma_y)}$$

Compares the boosted electric field with the QED critical field, often the symbol  $\Upsilon=\chi$  or  $\Upsilon=3\chi/2$  is used

( $\epsilon$ ,  $p^{\mu}$ : electron energy/four-momentum,  $r_e \approx 2.8 \times 10^{-13} \,\mathrm{cm}$ 

*N*: particles per bunch,  $\sigma_{x,y,z}$ : r.m.s. bunch dimensions)

Faci	lity	Energy [GeV]	#Particles [10 <sup>10</sup> ]	$\sigma_x[\mu m]$	$\sigma_y[\mu m]$	$\sigma_z[\mu \mathrm{m}]$	$\chi_{\max}$
шс	base	250	2	0 474	0 0059	300	0.15
ile	upgrade	500	2	0.474	0.0000	500	0.30
CLIC	base	190	0.37	0.045	0 0000	44	1.5
CLIC	upgrade	1500	0.57	0.045	0.0009	44	12
	base	10	1.2	18	12	1.8	0.01
FACET II	upgrade	10	0.7	3	2	0.5	0.13
MACE	small- $\beta^*$	125	0.5	0.01	0.01	0.1	1300
	small-z		0.06			0.01	1700

K. Yokoya and P. Chen, Frontiers of Particle Beams, 415-445 (1992)

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F	acility	Energy [GeV]	#Particles [10 <sup>10</sup> ]	$\sigma_x[\mu\mathrm{m}]$	$\sigma_y[\mu m]$	$\sigma_z[\mu \mathrm{m}]$	$\chi_{\max}$
	base 250		2	0 474	0.0050	200	0.15
ILC	upgrade	500	2	0.474	0.0059	500	0.30
CLIC	base	190	0.27	0.045	0 0000	44	1.5
ULIC	upgrade	1500	0.57	0.045	0.0009	44	12
FACE	TII + <sup>Plasma</sup> lense	<b>a</b> 10	1	0.04	0.04	1	5
MACE	small- $\beta^*$	105	0.5	0.01	0.01	0.1	1300
MACL	small-z	125	0.06	0.01	0.01	0.01	1700

K. Yokoya and P. Chen, Frontiers of Particle Beams, 415-445 (1992)

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#### Synchrotron radiation: classical description



#### Synchrotron radiation: classical spectrum & total power



#### Characteristic scalings

- Small frequencies:  $dP/d\omega \sim (\omega/\omega_c)^{1/3}$
- Large frequencies:  $dP/d\omega \sim \sqrt{\omega/\omega_c} \exp(-\omega/\omega_c)$

Critical frequency:  $\hbar\omega_c = (2/3)\epsilon\chi$ Plot (left side):  $\chi = 10^{-3}$ , i.e. strong exponential suppression well before  $\hbar\omega = \epsilon$ 

#### **Total radiation power**

Power P (energy per unit time) emitted per electron:

$$\left| P \sim \alpha \cdot \frac{c}{l_{\rm f}} \cdot \hbar \omega_c \sim \alpha \cdot \frac{mc^2}{\hbar} \frac{\chi}{\gamma} \cdot \epsilon \chi = \alpha \chi^2 \frac{(mc^2)^2}{\hbar} \sim \frac{1}{m^4} \right|$$

exact:  $P = \alpha \chi^2 (2/3) m^2$ ;  $m^2 = 63.56 \times 10^6 \text{ W}$ ;  $\alpha \sim 1/\hbar$ ;  $\chi \sim \hbar$ Intuitive derivation: photon emission probability per formation time  $c/l_f$  is  $\alpha$ ; typical energy of the radiated photon:  $\hbar \omega_c$ 

Schwinger, On the Classical Radiation of Accelerated Electrons, Phys. Rev. **75**, 1912 (1949) Sebastian Meuren (Princeton University) Strong-field QED @ FACET-II

# The QED critical field in classical electrodynamics (CED)

- Photon emission and pair production: related by a crossing symmetry
- However, pair production has a kinematic threshold ( $\epsilon \geq mc^2$ )
  - $\longrightarrow$  Pair production is exponentially suppressed for  $\chi_\gamma \ll 1$
- Photon emission with  $\omega\gtrsim\omega_c$  shows similar suppression (tunneling exponent)
  - $\longrightarrow$  Violation of the uncertainty principle is exponentially penalized



- CED predicts a qualitative change at the critical field ( $\chi \gtrsim 1$ ):  $\rightarrow$  emission of photons with  $\hbar \omega > \epsilon$  feasible (wrong, CED breakdown)
- General conclusion: If a theory predicts a qualitative change at a certain scale, one should test this scale experimentally!

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# Why does radiation possess a classical limit?

Quantity	"Compensation" of quantum effects
$P = \frac{2}{3}\alpha\chi^2 \frac{(mc^2)^2}{\hbar}$	light quantization ( $\hbar\omega_c$ ) vs. emission probability ( $lpha$ )
$l_{f}=rac{\gamma}{\chi}rac{\hbar}{mc}$	Compton length $[\mathcal{X}_{\mathcal{C}} = \hbar/(\mathit{mc})]$ vs. critical field $(\chi)$

- In the classical limit (  $\chi \ll$  1)  $\hbar$  must disappear from all quantities
- Therefore, the formation length has to be macroscopic  $(l_f \gg \gamma \lambda_C)$ 
  - $\rightarrow$  Possible, as typical photon energy  $\hbar\omega_c = (2/3)\epsilon\chi \ll \epsilon$  is very small, therefore, the uncertainty principle can be satisfied:

#### Momentum conservation

$$\begin{pmatrix} \epsilon \\ p_x \\ 0 \\ 0 \end{pmatrix} + \begin{pmatrix} 0 \\ q_x \\ q_y \\ 0 \end{pmatrix} = \begin{pmatrix} \epsilon' \\ p'_x \\ p'_y \\ p'_z \end{pmatrix} + \begin{pmatrix} \omega \\ k_x \\ k_y \\ k_z \end{pmatrix}$$

 $\epsilon$  (**p**),  $\epsilon'$  (**p**'),  $\omega$  (**k**): energy (momentum) of the initial electron, final electron, emitted photon; **q**: momentum transfered by the field

#### **Uncertainty principle**

- After some algebra:  $|q_x| \sim \chi mc/\gamma$   $[\epsilon' = \epsilon - \omega \longrightarrow p'_x \approx p_x - k_x - \omega/(2\gamma^2)$  $\longrightarrow -q_x = \omega/(2\gamma^2) \sim \omega_c/\gamma^2 \sim \chi m/\gamma]$
- $k_y, k_z \sim k_x/\gamma$  are subleading (1/ $\gamma$  cone)

$$|q_x| I_f \sim rac{\chi mc}{\gamma} rac{\hbar \gamma}{\chi mc} \sim \hbar$$

#### How does QED fix the problems of CED at the scale $\chi \gtrsim 1$ ?

- If  $\chi \gtrsim 1$  classical electrodynamics predicts  $\omega_c \sim \epsilon \chi \gtrsim \epsilon$  (not possible)
- The recoil ( $\hbar\omega$ ) induced by the emitted photon becomes important  $\rightarrow$  quantization of the photon field must be taken into account
- Semiclassical approach: classical trajectory + photon recoil at the vertex

$$\frac{dP}{d\omega} = \frac{dP}{du}\frac{du}{d\omega}, \quad \frac{dP}{du} = -\alpha m^2 \frac{u}{(1+u)^3} \left\{ \int_z^\infty dt \operatorname{Ai}(t) + \frac{\operatorname{Ai}'(z)}{z} \left[ 2 + \frac{u^2}{(1+u)} \right] \right\},$$

 $z = (u/\chi)^{2/3}$ ;  $u \approx \omega/(\epsilon - \omega)$ ;  $du/d\omega \approx (1 + u)^2/\epsilon$ ; Classical limit:  $u \approx \omega/\epsilon \sim \chi \ll 1$  ( $z \sim 1$ )



# Suppression of radiation in the quantum regime

classical scaling (	$\chi \ll 1)$ quantum scaling $(\chi\gtrsim 1)$
formation length $I_f = \frac{\gamma}{\chi} \frac{\hbar}{mc}$	$I_f=rac{\gamma}{\chi}rac{\hbar}{mc}(1+\chi/u)^{1/3}$
$\underline{\qquad } \text{critical frequency}  \hbar\omega_{c} = \frac{2}{3}\epsilon\chi; \ \left[ u \approx \right]$	$\frac{\omega}{\epsilon} \sim \chi \Big]  \hbar \epsilon \gtrsim \omega_c; \ \Big[ u \approx \frac{\omega}{(\epsilon - \omega)} \gtrsim 1 \Big]$
• The emitted photon energy no long	ger increases with $\chi$
• The formation length decreases slow	wer due to the factor $(1+\chi/u)^{1/3}$
Total emitted power	General result vs. asymptotics
$P = -\alpha P_0 \chi^2 \int_0^\infty dz  z \frac{4u^2 + 5u + 4}{2(1+u)^4} \operatorname{Ai'}(z),$ $\frac{P}{\alpha P_0} \approx \frac{2^5 \Gamma(2/3)}{3^5} (3\chi)^{2/3} \approx 0.37 \chi^{2/3} \ (\chi \gg 1)$ $u = \chi z^{3/2}, \ P_0 = (mc^2)^2 / \hbar$ $\frac{\text{Intuitive derivation}}{P \sim \alpha \cdot c/l_f \cdot \hbar \omega \sim \alpha \chi^{2/3} P_0}$ photon emission probability per formation time $c/l_f$ is $\alpha$ ; typical photon energy: $\hbar \omega \lesssim \epsilon$	$\begin{array}{c} 10^{2} \\ 10^{1} \\ 10^{0} \\ 10^{-1} \\ 10^{-2} \\ 10^{-3} \\ 10^{-4} \\ 10^{-5} \\ 10^{-6} \\ 10^{-3} \\ 10^{-2} \\ 10^{-1} \\ 10^{-1} \\ 10^{-1} \\ 10^{-1} \\ 10^{0} \\ 10^{1} \\ 10^{2} \\ 10^{3} \\ \chi \end{array}$ dashed lines: $\chi \ll 1$ and $\chi \gg 1$ asymptotics
Baier, Katkov, Strakhovenko: Electromagnetic Proce	sses at High Energies in Oriented Single Crystals (1998

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#### Full quantum regime: intuitive derivation of the scaling

	$\hbar\!\sim\!\chi\ll$ 1: independence of $\hbar$	$m^{-3} \sim \chi \gg 1$ : independence of $m$
Formation length	$I_{\rm f} = rac{\epsilon}{\chi} rac{\hbar c}{(mc^2)^2} \sim m$	$l_{ m f}=rac{\epsilon}{\chi^{2/3}}rac{\hbar c}{(mc^2)^2}\sim\hbar^{1/3}$
Power	$Ppprox 0.66lpha\chi^2rac{(mc^2)^2}{\hbar}\sim rac{1}{m^4}$	$Ppprox 0.37lpha\chi^{2/3}rac{(mc^2)^2}{\hbar}\sim rac{1}{\hbar^{4/3}}$

• Even if  $(\epsilon \gg mc^2)$ : physical observables depend on the rest energy  $mc^2 \longrightarrow$  Relativity (length contraction):  $l_f \sim \epsilon$ , independence of  $\hbar$ :  $l_f \sim 1/\chi \longrightarrow$  Photon-electron coupling strength:  $P \sim \alpha$ , independence of  $\hbar$ :  $P \sim \chi^2$ 

- If  $\chi \gtrsim 1$ :  $p^2 = m^2$  and  $(e^2 p F^2 p)^{1/3}$  change "mass hierarchy"  $\rightarrow$  The leading-order contribution should be independent of m
  - [note: we cannot neglect  $e^2 p F^2 p$  if  $\chi \ll 1$ , radiation requires field]

#### Approximations employed and their breakdown

- So far the only quantum correction is the recoil at the emission vertex
- Emission of virtual photons (radiative corrections) neglected
- Emission of multiple photons within one formation length neglected These assumptions seem to break down in the regime  $lpha\chi^{2/3}\gtrsim 1$

## Emission of virtual photons: radiative corrections



- Many conceptual questions related to  $\alpha \chi^{2/3} \gtrsim 1$  remain unsolved
- So far the regime  $\alpha \chi^{2/3} \gtrsim 1$  has been considered as very interesting but experimentally unaccessible, even in the far future

#### Now we have a realistic road map to this scale

- $\longrightarrow$  Theory: strong reason to revisit & extend existing calculations
- $\longrightarrow$  Experiment: very first access to strongly-coupled QED!

Ritus, Sov. Phys. JETP 30, 1181 (1970); SM and Di Piazza, PRL 107, 260401 (2011)

# From CED to the fully nontperturbative quantum regime

- In general, also the emission of virtual photons must be taken into account
- If  $\alpha\chi^{2/3}\gg 1$  the emission of virtual photons is no longer perturbative
  - $\longrightarrow$  Conjecture by Ritus & Narozhny:  $\alpha\chi^{2/3}$  is true expansion parameter



$$\bigotimes = \underbrace{\underbrace{\bigcap}_{\mathcal{O}(\alpha\chi^{2/3})}}_{\mathcal{O}(\alpha^{2}\chi^{4/3})?} + \underbrace{\underbrace{\bigcap}_{\mathcal{O}(\alpha^{2}\chi^{4/3})?}}_{\mathcal{O}(\alpha^{2}\chi^{4/3})?} + \cdots$$

Exact electron wave function (top), Mass operator (bottom)

#### Different regimes of strong-field QED:

 $1 \chi \ll 1 : classical regime$ 

Quantum effects are very small, pair production is exponentially suppressed

- $\begin{array}{ll} \textcircled{0.2cm}{2} & \chi \gtrsim 1, \alpha \chi^{2/3} \ll 1: \mbox{ quantum regime (FACET II)} \\ Recoil and pair production are important, but the radiation field is a perturbation \\ \end{array}$
- **(a)**  $\alpha \chi^{2/3} \gtrsim 1$ : **fully nonperturbative regime (100 GeV collider)** *Perturbative treatment of the radiation field breaks down*

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# Emission of multiple photons: radiation reaction



- Each emitted photon carries only a very small fraction of the electron energy
- The electron energy is changed *adiabatically* over many emissions

## Quantum radiation reaction ( $\chi \gtrsim 1$ )

- The recoil of a *single* photon changes energy and trajectory significantly
- The changed electron trajectory strongly modifies the subsequent emissions A. Di Piazza et al., Rev. Mod. Phys. **84**, 1177 (2012)

# Summary: four options to study strong-field QED at SLAC



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# Thank you for your attention and your questions!