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## Acceleration in Crystals & CNTs : Ideas, Challenges, Opportunities

Summary of the Workshop on Acceleration In Crystals and Nanostructures (June 24-25, 2019; Fermilab)

Vladimir SHILTSEV (Fermilab), with input from T.Tajima, A.Sahai FACET-II Science Workshop, Oct 31, 2019

The goal of the **Workshop** was to assess the progress of the concept over the past two decades and to discuss key issues toward proof-ofprinciple demonstrations and next steps in theory, modeling and experiment.

The Workshop was endorsed by the APS DPB and GPAP, the International Committee on Ultra-High Intensity Lasers (ICUIL) and the International Committee on Future Accelerators Panel on Advanced and Novel Accelerators (ICFA ANA). Fermilab, June 24-25, 2019

Workshop on Beam Acceleration in Crystals and Nanostructures

https://indico.fnal.gov/event/19478/

Organized by T. Tajima (UCI) and V. Shiltsev (FNAL) Proc.Eds.: S.Chattopadhyay, G. Mourou, V. Shiltsev, T. Tajima Endorsed by: APS GPAP & DPB, ICFA ANA, ICUIL, NIU







The concept of beam acceleration in solid-state plasma of crystals or nanostructures like CNTs has the promise of ultra-high accelerating gradients O(1-10) TeV/m, continuous focusing and small emittances of, e.g., muon beams and, thus, may be of interest for future high energy physics colliders. The main objective of the Workshop is to assess the progress of the concept over the past two decades and discuss the key issues toward proof-of-principle demonstration and next steps in theory, modeling and experiment.



40 participants - 2 days - 22 presentations

#ChicagoML

### **Main Topics of Presentations and Discussions**

1. **overview** of the past and present theoretical developments toward crystal acceleration, ultimate possibilities of the concept;

- 2. concepts and prospects of **PeV colliders for HEP**;
- 3. effective crystal wake drivers: beams, lasers, other;
- 4. beam **dynamics** in crystal acceleration;
- 5. instabilities in crystal acceleration (filamentation, etc.);
- 6. acceleration in **nanostructures** (CNTs, etc);
- 7. muon sources for crystal acceleration;
- 8. application of crystal accelerators (X-ray sources, etc.);
- 9. astrophysical evidence of wakefield acceleration processes;

10. steps toward "proof-of-principle": **1 GeV over 1 mm**, open theory questions, modeling and simulations;

11. possible experiments at **FACET-II, FAST, AWAKE**, AWA, RHIC, LHC, CEBAF, or elsewhere

# Workshop Proceedings

 co-edited by Profs.
 Gerard Mourou (Ecole Polytech, 2018 Nobel Prize in Physics), Toshiki Tajima (UCI), Swapan Chattapdhyay (NIU) and Vladimir Shiltsev (Fermilab). Proceedings of the Workshop on

Beam Acceleration in Crystals and Nanostructures

(Fermilab, June 24-25, 2019)

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- Aakash Sahai (U.Colorado), Toshi (Fermilab)
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#### Solid-state Tube Wakefield Accelerator Using Surface Waves in Crystals 10. Carlo Maria Lazzarini (ELI Beamlines), et al

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# **Plasma Waves**



## Acceleration in Continuous Focusing Channel

T.Tajima, M.Cavenago, Crystal X-ray Accelerator, Phys. Rev. Lett., 59(13), 1440 (1987).





10<sup>22</sup> cm<sup>-3</sup> → 10 TV/m,  $\lambda_p$ ~0.3µm 10<sup>24</sup> cm<sup>-3</sup> → 100 TV/m,  $\lambda_p$ ~0.03µm

Synchtrotron radiation losses balance energy gain: 0.3TeV for positrons 10 000 TeV for muons (+) 1000 000 TeV for protons

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Chen P, Noble R J AIP Conf. Proc. 398 273 (1997)

# Linear µ+µ- Crystal X-ray Collider



### **Basic ("Level 0") Questions**

- 1. Structured media:
  - crystals
  - carbon nanotubes
- 2. Drivers of wakefields in crystals:
  - lasers
  - beams
  - else
- 3. Particles to accelerate:
  - muons
  - electrons
  - protons



# **Crystals**

## Used in RHIC, Tevatron, LHC for *p,i* collimation (V.Shiltsev)

At SLAC for channeling and bending *e+/e-*(U.Wienands)

# Possibility of nanomodulation (G.Stupakov)



**Fermilab** 

## (Carbon) Nanotubes - CNT

Ends usually closed

Y.M.Shin, et al, Nuclear Instruments and Methods in Physics Research B 355 (2015) 94–100

#### - Carbon Nano Tubes (CNTs)

#### Rolling single graphene-sheet





Entrance to a rope of (10,10) SWNTs

~14 Å

#### Front view of (110) channels in Si crystal



- Wider channels: weaker de-channeling
- Broader beams (using nanotube ropes)
- Wider acceptance angles (< 0.1 rad)
- Lower minimum ion energies (< 100 eV)
- 3-D control of beam bending over
- greater lengths



Lattice Constant of Unit Cell ~ 4 A

Lattice Constant of Unit Cell ~ 14 A



# Y.M.Shin *et al*: CNT Experiment at FAST Slit-mask micro-bunching

**1 nC;**  $\lambda_{mb}$  = 100 μm



# Other Ideas: Combine Channels (funnel)





premodulated beam

self-modulating instability

## Wakes by Lasers – New Laser Concepts

#### **Chirped-Pulse Amplification**

#### (3 talks)

Short pulse oscillator Dispersive delay line

Chirped-pulse amplification involves stretching the pulse before amplifying it, and then compressing it later

We can stretch the pulse by 10,000, amplify it, and then r

#### 7 fibers coherently Added (2019)

**HOW THE ICAN LASER AMPLIFIER WORKS** 

The International Coherent Amplifier Network laser will produce high peak power and high average power, mitigating heat.

> An oscillator produces a 30-fs pulse.

> > 4

a single-mode optical fiber amplifier and passes through a pair of diffraction gratings, The which stretch it by pul around 105 times at the mu mJ level. The stretching mo separates the pulse's Eac components, producing the a "rainbow in time." mJ

The pulse is fed into

The stretched pulse is coupled to a multiplicity of singlemode fiber amplifiers. Each fiber will amplify the input pulse to the mJ level.

> The same operation is repeated in a second and third amplifier stage, where each fiber amplifier of the first stage feeds a multiplicity of single-mode amplifiers. In turn, each fiber will amplify its input to the mJ level.

The phase of each pulse is preserved in each amplifier. Finally, the chirped pulses are combined and phased. They now form single pulses and are compressed by a pair of gratings. The pulse energy can now be tens of joules; duration corresponds to initial pulse duration.

After exiting the last amplifier, the pulse is focused by a spherical or paraboloidal mirror.

The resulting pulse is short (30 fs), but the energy is enormous (30 J).

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30 fs - 30 nJ

30 fs - 30 J

## **Thin-Film Compression**



### Wakes by lons

• Bethe formula for ionization energy loss:

$$-\left\langle rac{dE}{dx}
ight
angle = rac{4\pi}{m_ec^2}\cdot rac{nz^2}{eta^2}\cdot \left(rac{e^2}{4\piarepsilon_0}
ight)^2 \cdot \left[\ln\!\left(rac{2m_ec^2eta^2}{I\cdot(1-eta^2)}
ight) - eta^2
ight]$$

• For high Z (of ions passing thru):

$$E_i \approx 2[\text{MeV}/(\text{g/cm}^2)] \times Z^2$$

• ie ~1 TV/m for Z=70-80 in silicon. Naturally, one can envision these ions either channeling in crystals ahead of the accelerating particles (e.g., muons) or being well aligned with them so the latter are always kept in sync with accelerating wake. At present, the highest energy heavy ions are available at RHIC (100 GeV/u gold, Z=79) and LHC (2.5 TeV/u lead, Z=82) and the dephasing length  $2\gamma_p^2\lambda_p$  can be as long as few cm - few meters.

# Visit to AWAKE (Oct. 2019... *ions* from SPS and *beam diagnostics* available)





## **Pre-Modulated Beam Possibilities (G.Stupakov)**

#### LCLS undulator at SLAC

This is achieved in a  $\sim$  100 m long magnetic undulator through the mechanism of an FEL instability.



#### Microbunched electron beam

FEL radiation is produced by an electron beam that is microbunched with the period  $\lambda_r.$ 



## Self-modulated long bunches AWAKE

ξ (mm)





ξ (mm)

20

ETTER

OPEN 10.1038/s41586-018-0485-4

#### Self-Modulation Instability in AWAKE p+ Bunch



Comparison of the proton-bunch longitudinal profile (left, no plasma) with the profile for a bunch passing through plasma (right), showing the strong modulation of the bunch.

#### ACCELERATOR TECHNOLOGY AWAKE makes waves



Beam at



- Compression XYZ8x7x2 um , 2 nC  $\rightarrow$ 
  - n\_e~0.6e19 cm<sup>-3</sup>
- Compression X Y Z 2x2x0.4 um , 2 nC  $\rightarrow$  n\_e~2e20 cm<sup>-3</sup>
- Peak currents: 70...100...300 kA !
- What can be studied there:
  - Weibel(filamentation) and SMI instabilities effect of structured media (?)
  - Muon production and channeling
  - Acceleration

#### **Experimental Study of Current Filamentation Instability**

B. Allen,<sup>1,\*</sup> V. Yakimenko,<sup>2</sup> M. Babzien,<sup>2</sup> M. Fedurin,<sup>2</sup> K. Kusche,<sup>2</sup> and P. Muggli<sup>3,1</sup>

<sup>1</sup>University of Southern California, Los Angeles, California 90089, USA <sup>2</sup>Brookhaven National Laboratory, Upton, New York 11973, USA <sup>3</sup>Max Planck Institute for Physics, Munich, Germany (Received 2 July 2012; published 2 November 2012)

Current filamentation instability is observed and studied in a laboratory environment with a 60 MeV electron beam and a plasma capillary discharge. Multiple filaments are observed and imaged transversely at the plasma exit with optical transition radiation. By varying the plasma density the transition between single and multiple filaments is found to be  $k_p \sigma_r \sim 2.2$ . Scaling of the transverse filament size with the plasma skin depth is predicted in theory and observed over a range of plasma densities. Lowering the bunch charge, and thus the bunch density, suppresses the instability.



## **FACET-II**

- Experiment E305: Beam filamentation and bright gamma-ray bursts
  - Sébastien Corde(ÉcolePolytechnique/LOA)
  - Ken Marsh (UCLA)
  - Frederico Fiuza (SLAC)



## "Prepare for FACET-II" : what can be done

- Simulations; define optimal configuration beams, CMTs vs crystals, etc
- Hardware assembly and tests, eg at FAST
- Beam pre-test at FAST (50-300 MeV ILC type beams)
  - E.g., *e* channeling and SMI in the CNTs
- Consider muon production, capture and channeling, acceleration
  - Then it can be expanded to FACT-II (10 GeV *e* to mu) or else where (BELLA has GeV *e*- beams)

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• Simplified schemes (for FAST/FACETII)



crystals &

### **Toward Experimental Studies of "Xtal Acc"**

Wakefield **Muon** Accelerator: ultrashort, micron-scale µ<sup>±</sup> beams

A. Sahai (CU), V. Shiltsev (Fermi), T. Tajima (UCI)

Proposal

1. µ<sup>±</sup> production & characterization – ultrashort e<sup>-</sup> beam driven

(in future, compact laser-driven e<sup>±</sup> -beam [when stable beams avail.]) We need to better understand the muon spectra from the foil as a majority of the muons are not relativistic. Nobody studied the muon spectra from an ultrashort 10um waist-size multi-GeV electron beam before

#### e<sup>±</sup> -beam driven wakefield μ<sup>±</sup> acc. – can go along with foil-e<sup>+</sup> proposal (extra μ<sup>±</sup> detectors, foils are quite similar due to same scaling)

The e-beam wakefield will only trap a very small percentage (already relativistic) of the 10,00 muons produced. Secondly, the electron beam exiting the foil may be degraded and may not drive a nice wake.

 Iaser driven wakefield μ<sup>±</sup> acc. – stage-2 is laser-driven (better control over wake velocity, more trapping by velocity matching)





#### (A.Sahai)

#### e<sup>-</sup> beam wakefield **Muon** Accelerator

- e<sup>-</sup> beam driven wakefield phase velocity very HIGH (gamma > 10000) near the speed of light
- only efficiently trap the µ<sup>±</sup> that are already relativistic produced in the external foil or in early stages of the plasma
- e-beam wakefield will only trap a small fraction (already relativistic) of the 10,000 µ<sup>±</sup> pairs
- Moreover, e<sup>-</sup> beam exiting the foil may be degraded may not drive a high-quality wake

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under further investigation !



## (A.Sahai)

PHYSICAL REVIEW SPECIAL TOPICS - ACCELERATORS AND BEAMS 12, 111301 (2009)

#### ultrashort, micron-scale $\mu^{\pm}$ production & characterization

#### μ<sup>±</sup> production rate



### (A.Sahai)



laser group velocity = wakefield phase velocity

laser group velocity ~ controlled using plasma density

wakefield phase velocity can be tuned to match µ<sup>±</sup> spectra & optimize trapping in the wakefield

**FACET-II beam** produces  $\mu^{\pm}$  in 1 cm thick target.

**FACET-II laser** driven wakefield accelerator to trap and accelerate  $\mu^{\pm}$ 

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#### **Connection to Cosmos - Toshi Ebisuzaki(RIKEN)**

EM Waves

# Formation of extragalactic jets from black hole accretion disk

Extragalactic

jet -

EM W7

Jes

Growing interest in the multi-messenger astronomy community – see eg S.Lukin (NSF) talk at the GPP/GPAP meeting (Oct. 22, 2019)

Excess map: three hot spots Aab et al. (2018) Astrophys. J. Letters, 853, L29

Model Flux Map - Starburst galaxies - E > 39 EeV





## Summary

- Acceleration of  $\mu$ 's in crystals or CNTs has great promise
- There are many issues related to *muon production, channeling* and acceleration
- Some modes of the crystal/CNT excitations can be tested at beam facilities such FACET-II, FAST, BELLA, AWAKE, etc
- Beam filamentation is of serious concern and can be studied first at, e.g., FAST, then elsewhere (FACET-II, etc)
  - Past experience and hardware very helpful
- Also can be tried at FACET-II : i) muon production; ii) muon detection; iii) experiment integration; iv) calibration of models
- A lot of serious work ahead to understand the most optimal ways to excite Xtals/CNTs, explore beam acceleration and dynamics via theory, modeling, and experiment



Thank You for Your Attention!



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Several interesting proposals for further explorations or experimental tests were made by Sahel Hakimi, et al. (University of California, Irvine, on how to drive wakes in CNTs by ultimate or existing X-ray pulses from, e.g., the LCLS SASE FEL); by Aakash Sahai, et al. (University of Colorado, on production of detectable number of muons and their subsequent acceleration either at BELLA or FACET-II facilities); by Vladimir Shiltsev, et al. (Fermilab, on demonstration of effective micromodulation of electron beams at FAST and FACET-II and subsequent experiments with micromodulated beams sent through CNTs at FAST with kA peak current type beams and then at the FACET-II facility with upto 300 kA bunches, e.g., to demonstrate the CNT channeling or to study the electron beam filamentation phenomena in structured materials); by Gennady Stupakov (SLAC, on possibility to use 1nm-SASE-modulated electron bunches at the end of LCLS-I undulators to excite crystals and demonstrate acceleration); by Johnathan Wheeler, et al., (Ecole Polytechnique, to use the APOLLO laser facility to demonstrate Peta-Watt optical pulses/single cycle pulses via thin-film-compression technique); by Valery Lebedev (FNAL, to explore effectiveness of the wake excitation in crystals or CNTs by high-Z high energy ions, e.g. by 450 GeV ion beams from the CERN SPS available at the AWAKE facility, and observation of possible acceleration of externally injected electrons).



#### Nanotubes (2)





## **Collider considerations**

$$\mathrm{d}N/\mathrm{d}t = -N/\gamma\tau_0 \quad \frac{N}{N_0} \approx \left(\frac{m_\mu c^2}{E}\right)^\kappa$$

 $A \sim 1 \ {
m \AA}^2 = 10^{-16} \ {
m cm}^2 \ N_0 \sim 10^3$  particles

 $L = fN^2/A = f \times 10^{16} \times 10^6 n_{\rm ch} \ [\rm cm^{-2} \ s^{-1}]$ 

milab

 $\kappa = (m_{\mu} c / \tau_0 G) \ll$ 

 $1/\ln(E/m_{\rm u}c^2)$ 

$$L [\mathrm{sm}^{-2} \,\mathrm{s}^{-1}] \approx 4 \times 10^{33-35} \frac{P^2 \,[\mathrm{MW}]}{E^2 \,[\mathrm{TeV}] \, fn_{\mathrm{ch}}[10^8 \,\mathrm{Hz}]}$$

#### Table 4. Options for future particle colliders.

Collider type	Dielectric based	Plasma based	Crystal channeling
Accelerating media	Microstructures	Ionized plasma	Solid crystals
Energy source: option 1 option 2	Optical laser e <sup>-</sup> bunch	e <sup>–</sup> bunch Optical laser	X-ray laser
Preferred particles	Any stable	e <sup>-</sup> , μ <sup>-</sup>	$\mu^+$ , $p^+$
Max accelerating gradient, $\text{GeV}\text{m}^{-1}$	1-3	30-100	$100 - 10^4$
CM energy reach in 10 km	3-10	3-50	$10^3 - 10^5$
Number of stages/10 km: option 1 36 option 2	$\frac{10^5 - 10^6}{10^4 - 10^5} $ 6	6/2 <u>46</u> 2019₄	Shlltsev   2019 FACET-II Science Workshop

### Ultimate Testbed FACET-II | Facility for Advanced Accelerator Experimental Tests



Electron Beam Parameter	Baseline Design	Operational Ranges	Positron Beam Parameter	Baseline Design	Operational Ranges
Final Energy [GeV]	10	4.0-13.5	Final Energy [GeV]	10	4.0-13.5
Charge per pulse [nC]	2	0.7-5	Charge per pulse [nC]	1	0.7-2
Repetition Rate [Hz]	30	1-30	Repetition Rate [Hz]	5	1-5
Norm. Emittance γε <sub>x,y</sub> at S19 [μm]	4.4, 3.2	3-6	Norm. Emittance γε <sub>x,y</sub> at S19	10, 10	6-20
Spot Size at IP $\sigma_{x,y}$ [µm]	1 <b>8</b> , 12	5-20	Spot Size at IP $\sigma_{x,y}$ [µm]	16, 16	5-20
Min. Bunch Length σ <sub>z</sub> (rms) [μm]	1.8	0.7-20	Min. Bunch Length $\sigma_z$ (rms)	16	8
Max. Peak current I <sub>pk</sub> [kA]	72	10-200	Max. Peak current Ipk [kA]	6	12

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#### **High Energy Particle Physics: Progress and Challenges**

• Collider physics – dominated by the LHC til 2038



Neutrino physics – multi-mW beams at Fermilab ~ till ~2040
 Fermilab

## What's Next?

#### • HEP Int'l Community planning:

- European Strategy 2012-2013
- US "Snowmass", P5, HEPAP 2013-2014
- European Strategy Update 2019-2020
- ILC250 in Japan? Decision by Feb. 2020
- Potential of CepC in China?
- US "Snowmass", P5, HEPAP 2019-2022
- Planning for longer term HEP future in general (20-50 yrs) and "Post-2026" Era in particular (next cycle)
  - Collider Physics Step 1: Higgs Factory(ies)
  - Collider Physics Step 2: Energy Frontier (50-100 TeV pp or 6-15 TeV lepton)

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#### Challenges:

– a) Cost and Feasibility; b) What's in the "Far Future"?

## Glimplse onto "Cost and Feasibility"

	Project	Туре	Energy [TeV]	Int. <u>Lumi.</u> [a <sup>.1</sup> ]	Oper. Time [V]	Power [MW]	Cost
	ILC	ee	0.25	2	11	129 <mark>(upgr</mark> . 150-200)	4.8-5.3 GILCU + upgrade
			0.5	4	10	163 (204)	7.8 GILCU
			1.0			300	?
	CLIC	ee	0.38	1	8	168	5.9 GCHF
			1.5	2.5	7	(370)	+5.1 GCHF
			3	5	8	(590)	+7.3 GCHF
	CEPC	ee	0.091+0.16	16+2.6		149	5 G\$
			0.24	5.6	7	266	
	FCC- <u>ee</u>	ee	0.091+0.16	150+10	4+1	259	10.5 GCHF
			0.24	5	3	282	
 			0.365 (+0.35)	1.5 (+0.2)	4 (+1)	340	+1.1 GCHF
	LHeC	ep	60 / 7000 GeV	1	12	(+100)	1.75 GCHF
	FCC-hh	рр	100	30	25	580 (550)	17 GCHF (+7 GCHF)
	HE-LHC	рр	27	20	20		7.2 GCHF
_	6/24/	2019	From D.S	Schulte, EPF	PSU Symposi	ium, Granad	a, Spain,. May 2019

# A Vision for "Ultimate" Colliders

- Post-100 TeV "Energy Frontier" assumes
  - ✤ 300-1000 TeV (20-100 × LHC)
  - "decent luminosity" (TBD)

# • Surely we know: circular collider

For the same reason there
 is no circular *e+e-* collider above
 Higgs-F there will be no circular *pp* colliders beyond 100 TeV → LINEAR

2. Electrons radiate 100% linear collider beam-strahlung (<3 TeV) and in focusing channel  $L \propto (<10 \text{ TeV}) \rightarrow \mu + \mu$ - or pp

 $\eta_{\text{linac}} P_{wall} N_{\gamma}$ 

 $L \propto \frac{\eta P_{wall}}{E^3} \frac{\xi y}{\beta_w}$ 

6/24/2019

# "Phase-Space" is Further Limited

- "Live within our means": for 20-100×LHC
   < 10 B\$</p>
  - **☆** < 10 km
  - < 10 MW (beam power, ~100MW total)</p>
- →New technology should provide >30 GeV/m @ total component cost <1M\$/m (~NC magnets now)

SC magnets equiv. ~ 0.5 GeV per meter (LHC)

3. Only one option for >30 GeV/m known now: <u>dense plasma</u>→ that excludes *protons*→ <u>only *muons*</u>



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## Paradigm Shift : Energy vs Luminosity



## What Do We Know about Crystals?



- Strong inter-planar electric fields ~10V/A=1GV/cm
- Very stable, can be used for
  - deflection/bending (works)
  - focusing (works)
  - > acceleration (*if excited*)

 $l_{\rm d}\,[{\rm m}] \sim E\,[{\rm TeV}]$ 

**T980 experiment** at **Tevatron**, N.Mokhov et al JINST 6 T08005 (2011)



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## **Bent Crystals in the 7 TeV LHC Beams**

6/24/2019







## 2015-2017 CRYSTAL CHANNELING EXPT @ FAST

• P.Piot, T.Sen, A.Halavanau, D.Edstrom, J,Hyun, et al



# What Do We Know About Acceleration in Xtals and Nanostructures (besides 1...10...100 tV/m)

- 1987 the idea: T.Tajima and M.Cavenago,
  - Bormann angle X-ray injection
  - Phys. Rev. Lett. 59 (1987), 1440
- 1990's P.Chen and R.Noble, scattering and cooling considerations, crystal damage, etc
  - SR losses balance E gain: 0.3TeV for e+, 10 PeV for mu+, 1000 PeV for p+
  - AIP Conf. Proc. 398 (1997), 273
- 2008 I.Dodin and N.Fisch, theory of acceleration in plasma channels, scattering, friction, damping
  - Phys. Plasmas 15 (2008), 103105
- 2012 V.Shiltsev, prospects of linear crystal muon colliders
  - Phys. Uspekhy 55 (2012), 965
- 2010's Prospects of superlasers, superbeams (FACET-II), CNTs
  - Shin, APL105(2014),114106; NIMA355(2015),94 ; Zhang,et al PRAB 19 (2016),101004

#### Ways to excite the crystal (1)



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### **Crystal Excitation by X-Rays**



Tajima, Cavenago, Phys. Rev. Lett. 59 (1987), 1440

FIG. 1. Bormann anomalous transmission. When the x rays are injected at the Bragg angle, the Bormann effect takes place. Particle beams are injected along the crystal axis.

Need 40keV high peak power x-rays

now available from SASE FELs like LCLS

Gradients >1GV/cm

- Muons preferred
  - No bremstrahlung, no nucl.
- $\mu$ + rad length 10^9 cm

#### Ways to excite the crystal (2)



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#### Ways to excite the crystal (3)



#### **Nanotubes**



#### Ways to excite the crystal (4)



# Ways to Excite Nanotubes/Crystals (5) by Optical SASE modulated electron beams

\* or slit-masking in low energy beams  $\frac{10^{22} \text{ cm}^{-3} \rightarrow \lambda_{p} \sim 0.3 \mu \text{m}}{10^{24} \text{ cm}^{-3} \rightarrow \lambda_{p} \sim 0.03 \mu \text{m}}$ 



#### Ways to Excite Nanotubes/Crystals (6) by Self-Modulation Instability in long(er) charge particle beams



#### SMI: Self-Modulation Instability (in 400 GeV protons)



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#### Secondary particle production 300 MeV electron on 2 radiation length of carbon target



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#### Positive pions has very big positron/proton escort, large angle & high momentum; negative pion could be focused and extracted





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#### More pion could be obtained for larger thickness & low Z. target material. Larger thickness – large radiation problem, low Z - longer target (more difficult to collect)



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# Low radiation scenario

#### Two production targets – photon production and pion/muon production.

Photon production target should be thin (~10% radiation length). Primary electrons can be swept and miss pion/muon target – compact muon source design (Nagamine at al 2001).

They made analytical estimate for 10% tungsten photon and 10 cm carbon pion/muon production target. For pion produced at 45 degree with acceptance 1 steradian and momentum from 150 to 163 MeV/c they got 3.5  $10^{-9} \pi$ /electron in 2001 (4.7  $10^{-8}$  π/electron in 2016). Our simulation shows that at 45 degree pions have heavy electron/positron escort, but at 90 degree we could get 3.7 10-8 negative pion/electron in above angular & momentum range. With such two target design we could get about 3 times more pion then with one 10% radiation length carbon target.

Large Omega muon optics channel could capture pion beam with dE=10 MeV and  $d\Omega$ =1 steradian and produce 0.4 muon/pion.



momentum (GeV/c)

# High Energy μ+μ- Colliders

## Advantages:

- µ's do not radiate / no beamstrahlung → acceleration in rings → low cost & great power efficiency
- ~x7 energy reach vs pp
- Offer "moderately conservative moderately innovative" path to cost affordable energy frontier colliders:



• ZDRs exist for 1.5 TeV, 3 TeV, 6 TeV and 14 TeV \* in the LHC tunnel

#### Key to success:

- Test facility to demonstrate performance implications muon production and 6D cooling, study LEMMA  $e^+-45$  GeV +  $e^-$  at rest  $\rightarrow \mu^+-\mu^-$ , design study
- <sup>62</sup> of acceleration, detector background and neutrino radiation



JINST Special Issue (MUON)

\* more like "strawman" parameter table

Input #120

#### **Current Filamentation Instability in Laser Wakefield Accelerators**

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### Weibel (Filamentation) Instability





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#### On the role of the purely transverse Weibel instability in fast ignitor scenarios

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The growth rate for the purely transverse Weibel instability is determined from relativistic kinetic theory using a waterbag distribution function in the momenta perpendicular to the main propagation direction of the beam. A parametric study is presented for conditions relevant to the fast ignitor. It is shown that for expected parameters the purely transverse Weibel instability will be significantly suppressed or even eliminated due to the transverse energy spread or emittance. © 2002 American Institute of Physics. [DOI: 10.1063/1.1476004]





$$\frac{\alpha}{\gamma_{b0}} \left( \frac{\beta_{x0}^2}{\beta_{z0}^2} + \frac{u_{x0}^2}{u_{x0}^2 + 1} \right) > \left( \frac{\overline{1}}{\gamma} \right) + \alpha \left( \frac{\overline{1}}{\gamma_b} \right).$$
(9)

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